

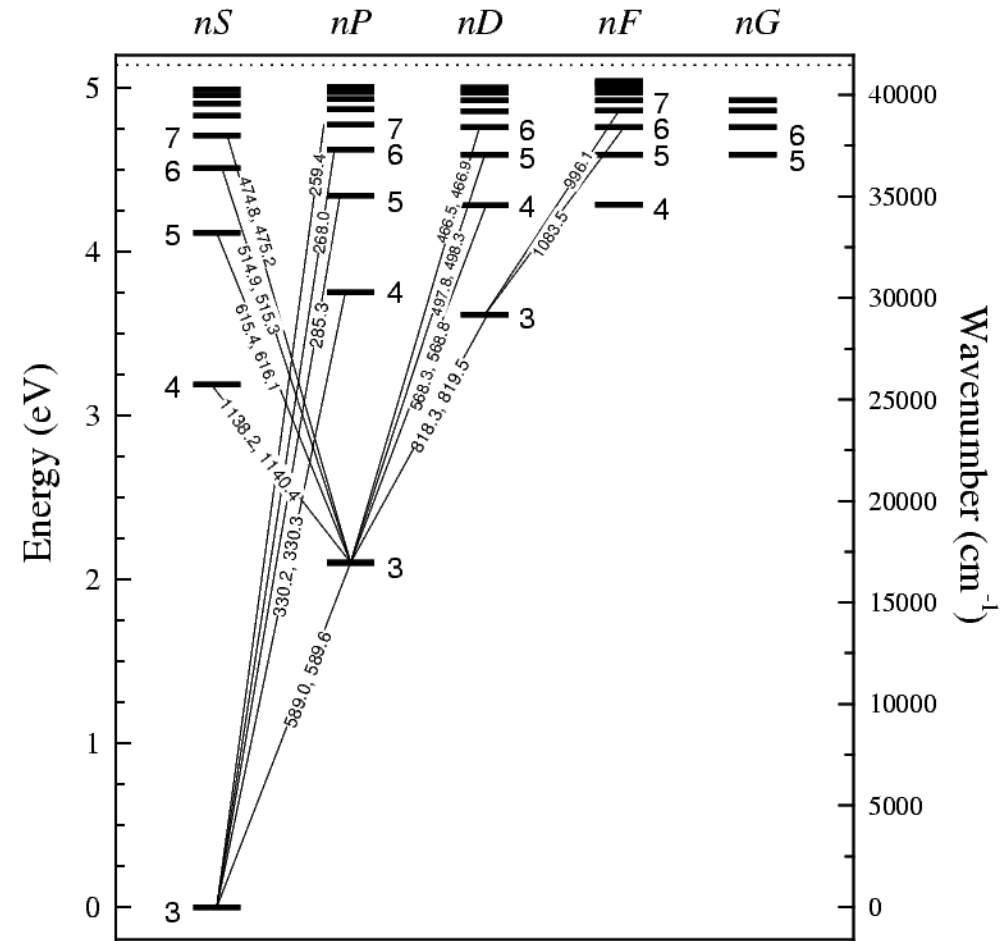
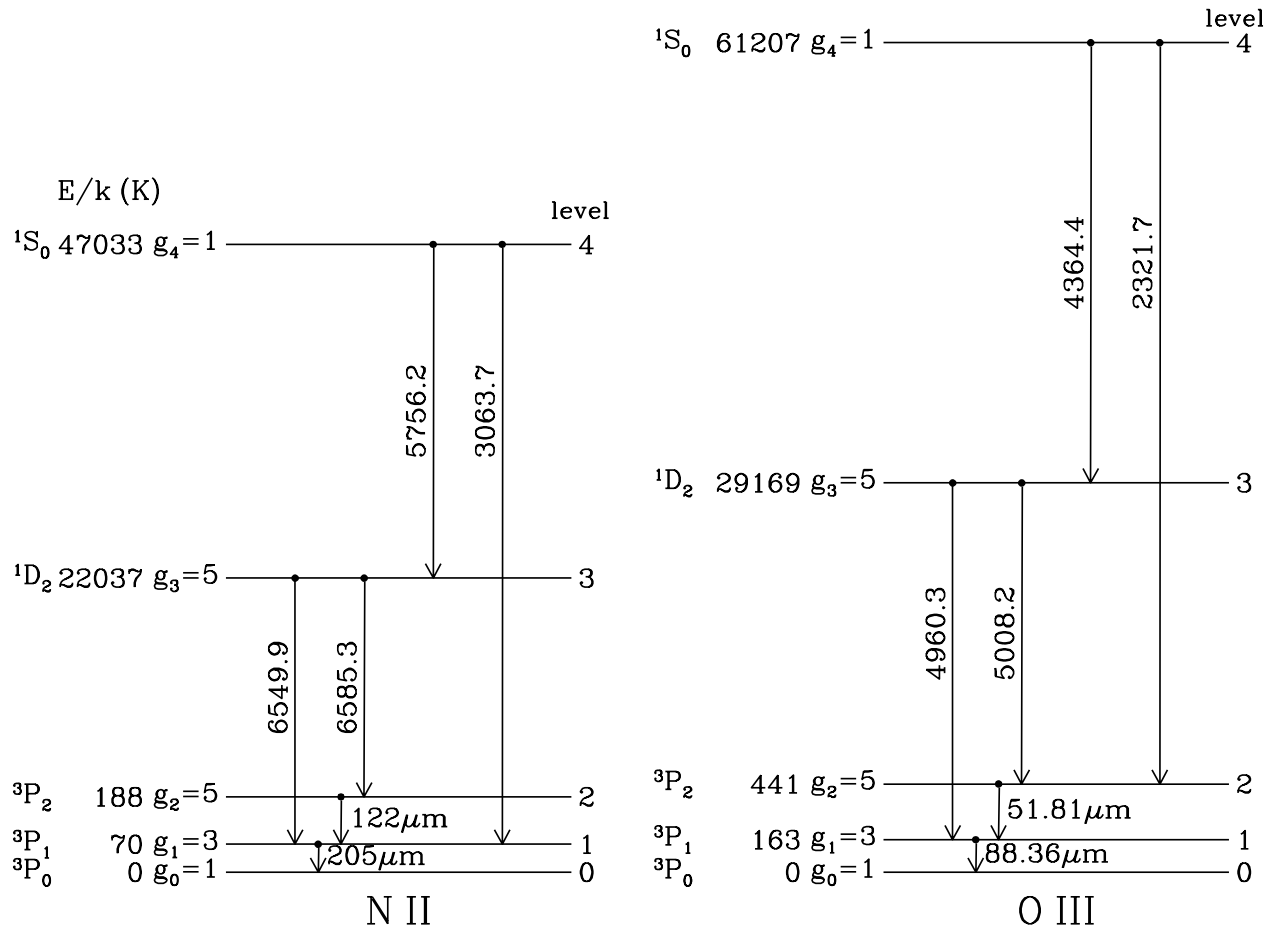
# INTERSTELLAR MEDIUM

- Stefano Bovino -



## Atomic Structure II

# Goal: Understanding these diagrams!



Z : 11

Ioniz. Pot. : 5.138 eV

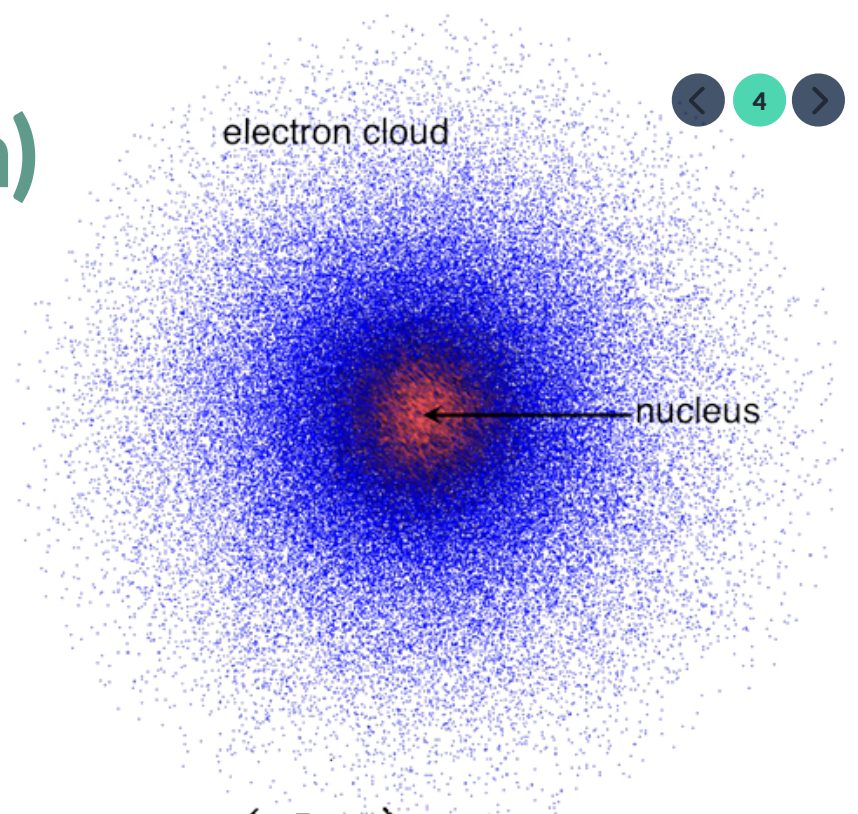
ground state :  $1s^2 2s^2 2p^6 3s$

# Atomic structure in a nutshell

- Electrons behave both as particles and waves (De Broglie)
- Quantisation of energy (Planck, Bohr)
- Wave function and energy of an atom (solving the S.E.)
- Wavefunction squared represents a probability (Born)

# Hydrogen atom (single electron)

electron cloud



nucleus

$$R_{nl}(r) = - \left\{ \frac{(n-l-1)!}{2n[(n+l)!]^3} \right\}^{1/2} \left( \frac{2}{na_0} \right)^{l+3/2} r^l e^{-r/na_0} L_{n+l}^{2l+1} \left( \frac{2r}{na_0} \right)$$

**Radial**

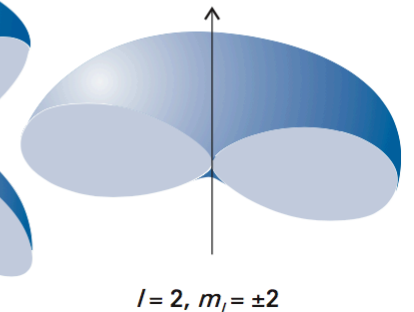
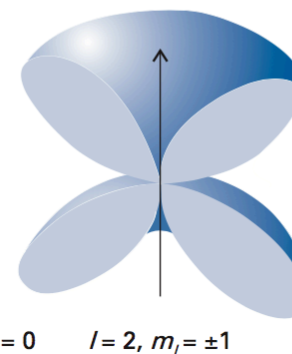
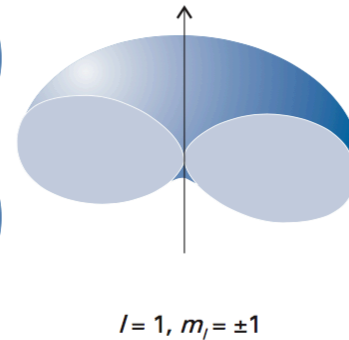
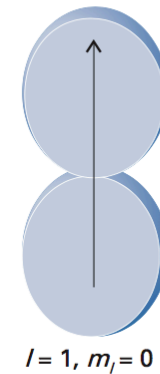
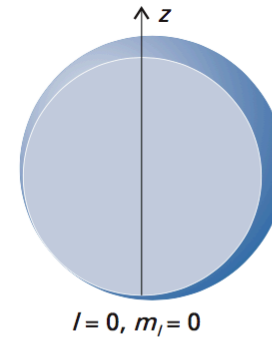
$$Y_l^m(\theta, \phi) = \left[ \frac{(2l+1)(l-|m|)!}{4\pi(l+|m|)!} \right]^{1/2} P_l^{|m|}(\cos \theta) \underbrace{e^{im\phi}}$$

**Angular**

# Quantum numbers describe orbitals

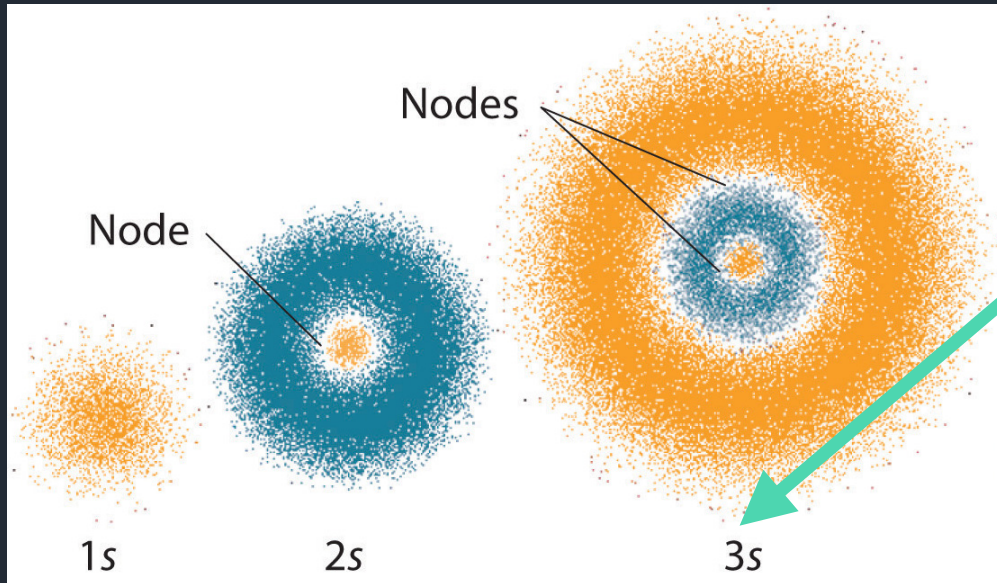
Orbitals represent regions of probability where an electron can be found

Each orbital can hold up to a maximum of 2 electrons (Pauli principle)

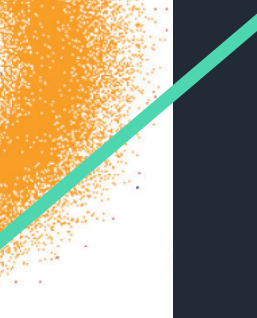



- $n$   $\longrightarrow$  principal quantum number
- Specifies the energy of the orbital and the shell
- Assume values  $n = 1, 2, 3\dots$
- Larger the value of  $n$  further away from the nucleus the electron is

Energy



Principal quantum number



- $l$   orbital (angular momentum) quantum number
- Defines the shape of the orbitals
- Assume values  $l = 0, 1, 2, n-1$

0

1

2

3

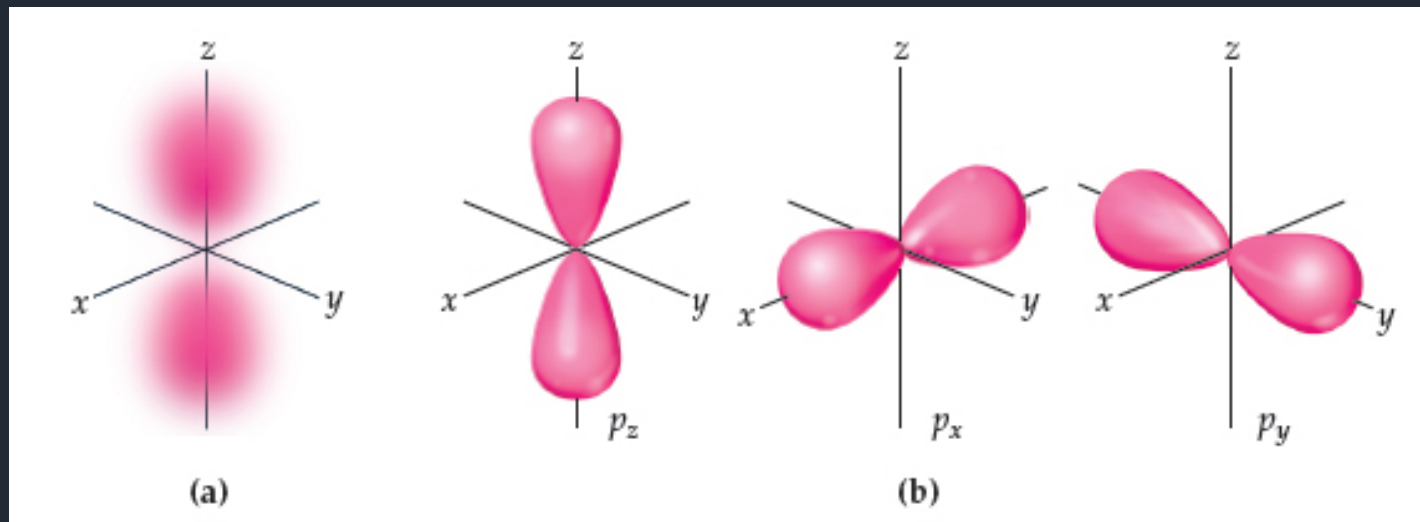
s-orbital

p-orbital

d-orbital

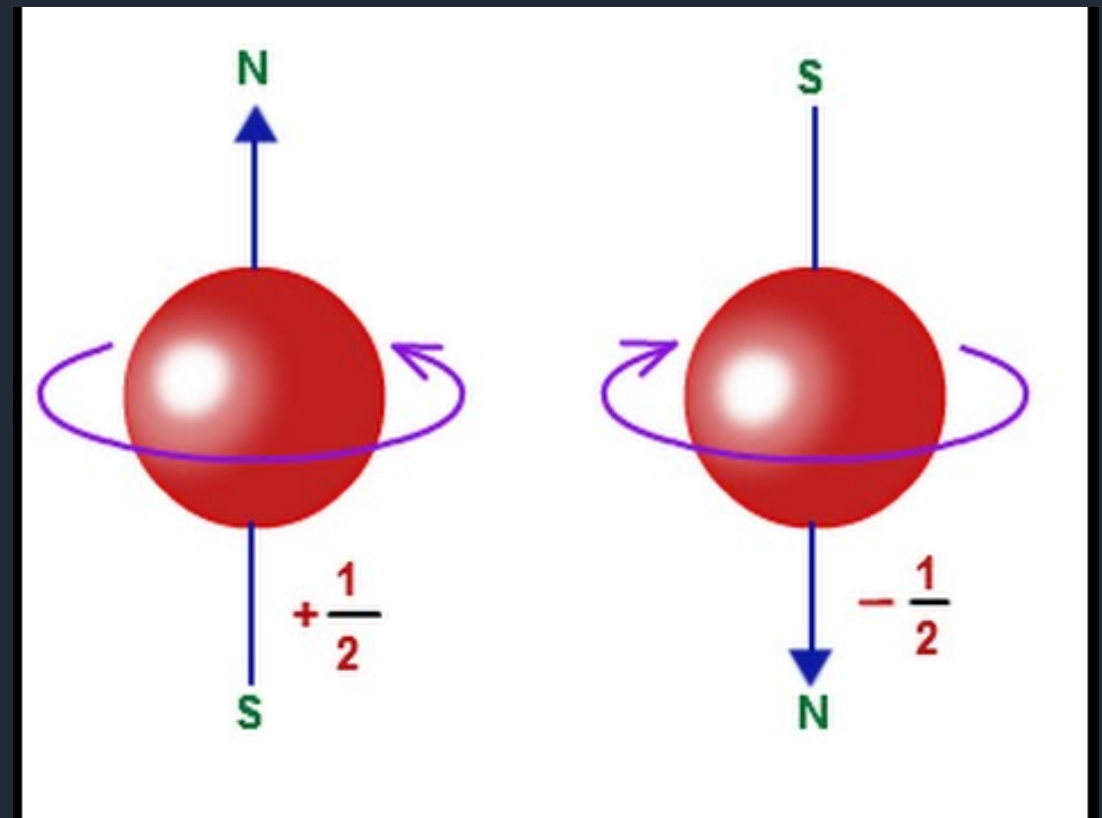
f-orbital

- $m_l$   $\longrightarrow$  magnetic quantum number
- Orientation of the angular momentum around the nucleus
- How many orbitals of that kind exist per energy level
- Assume values  $m_l = l, l-1, -l$
- $2l + 1$  values of  $m_l$  for a given value of  $l$





- $s, m_s$   $\longrightarrow$  spin magnetic quantum number (z-component)
- Intrinsic angular momentum of the electron
- Assume values  $m_s = -1/2, +1/2$



# Electronic configuration

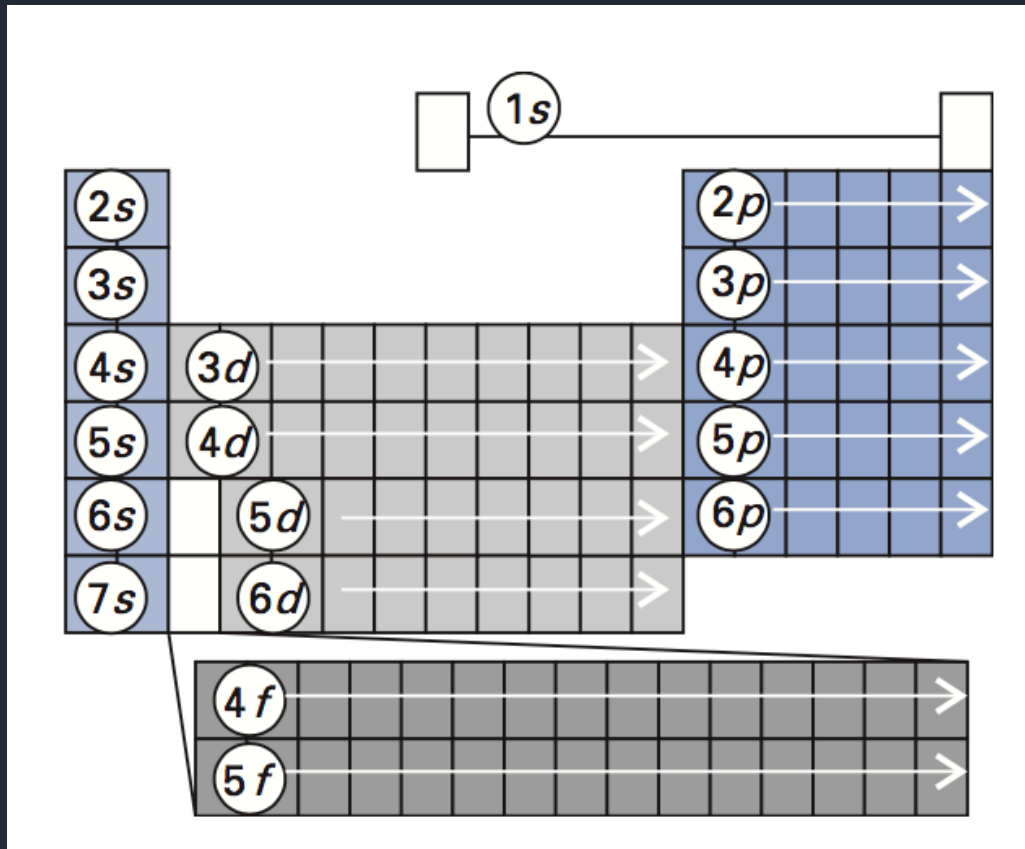
- **Configuration** specifies the orbitals that electrons occupy
- H ( $1s^1$ ) means one electron occupies a  $1s$ -orbital
- A single configuration (e.g.  $2p^1$ ) can split in more levels (different  $j$ )

**Hund's rules**

**Pauli principle**

# Multielectrons systems

## Aufbau principle



## Hund's rules

## Pauli principle

**Example: Cl or NII**

Six electrons



# Multielectrons systems

- Chemists (and spectroscopists) introduced the term symbols

multiplicity  $\rightarrow 2S+1$   $\{L\}$   $\leftarrow$  orbital angular momentum  
 $J \leftarrow$  Level

## Example: Hydrogen atom

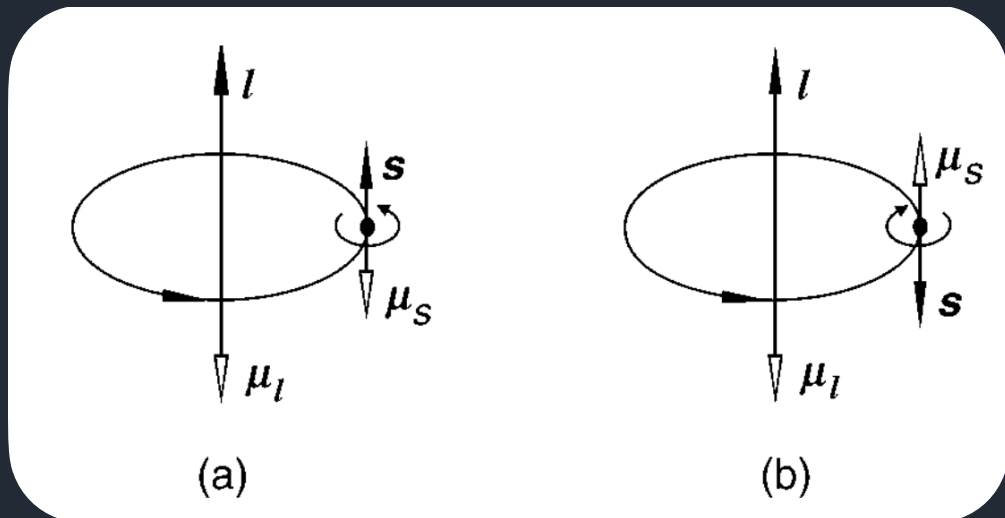
One electron

$1s^1$

What is the term symbol of the ground state?

# Spin-orbit coupling

- Rotation of an electrical charge generates a magnetic dipole (classical electrodynamics)
- The spin and orbital angular momenta (e-) couple
- This generates a splitting of a given energetic level (fine structure)
- This can be further split by the presence of a magnetic field (hyperfine structure)



The energy shift depends on L and S

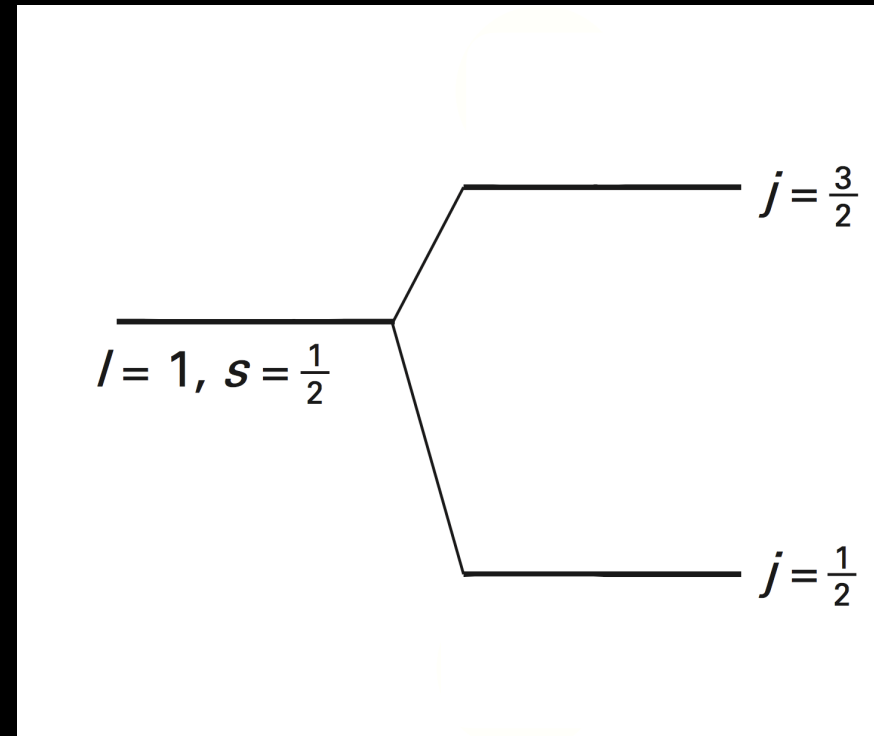
The left diagram shows an electron with spin  $S$  pointing up and orbital magnetic moment  $N$  pointing up. The spin magnetic moment  $\mu_s$  (green arrow) points up, and the orbital magnetic moment  $\mu_l$  (red arrow) points up. This configuration is labeled  $m_s = +\frac{1}{2}$ .

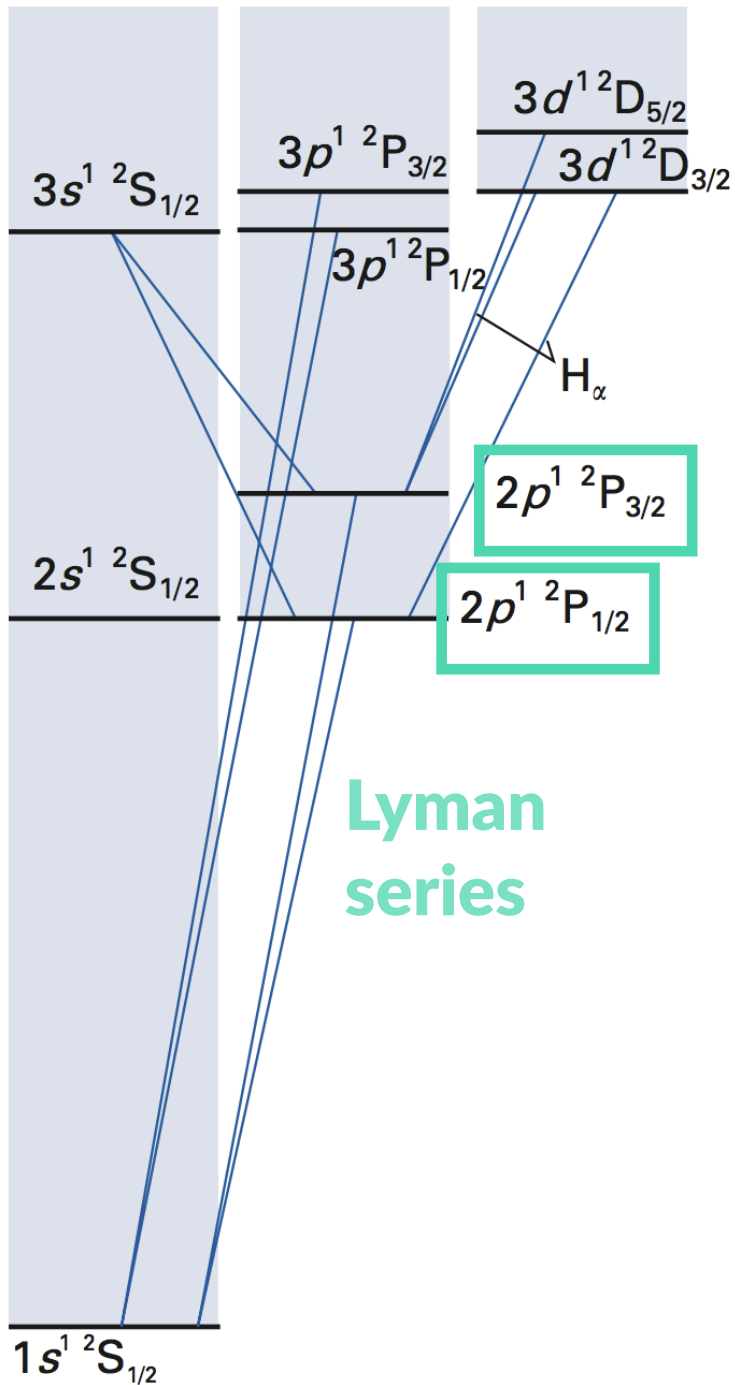
The right diagram shows an electron with spin  $S$  pointing down and orbital magnetic moment  $N$  pointing up. The spin magnetic moment  $\mu_s$  (green arrow) points down, and the orbital magnetic moment  $\mu_l$  (red arrow) points up. This configuration is labeled  $m_s = -\frac{1}{2}$ .

# Spin-orbit coupling (fine)

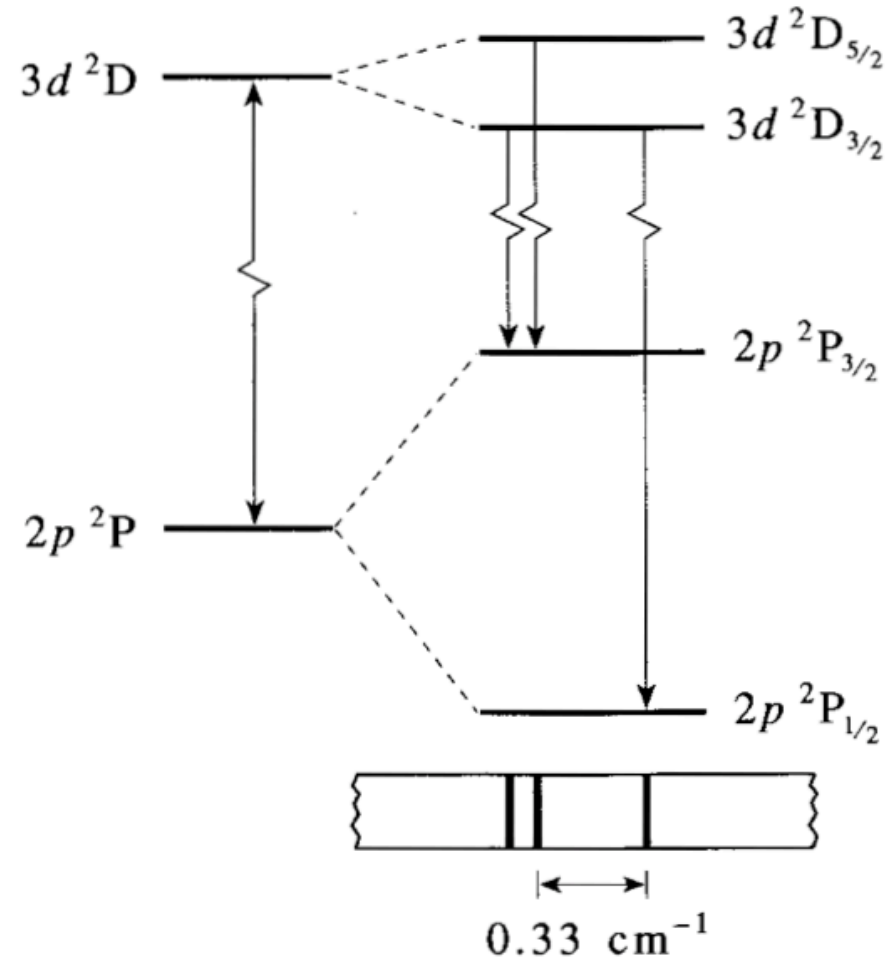
$$j = l + s$$

## Doublet



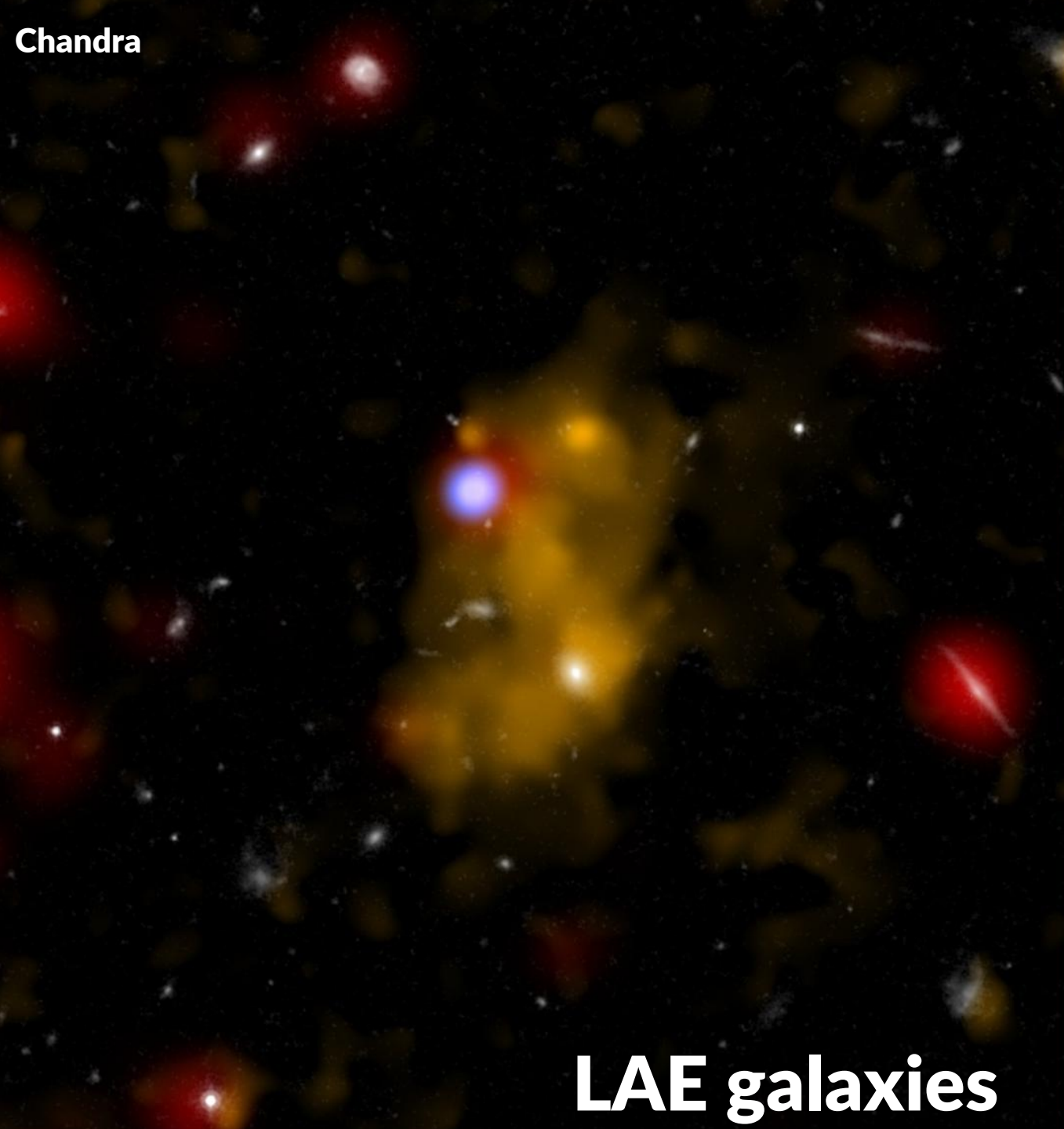


Lyman series



# Spin-orbit coupling

Chandra



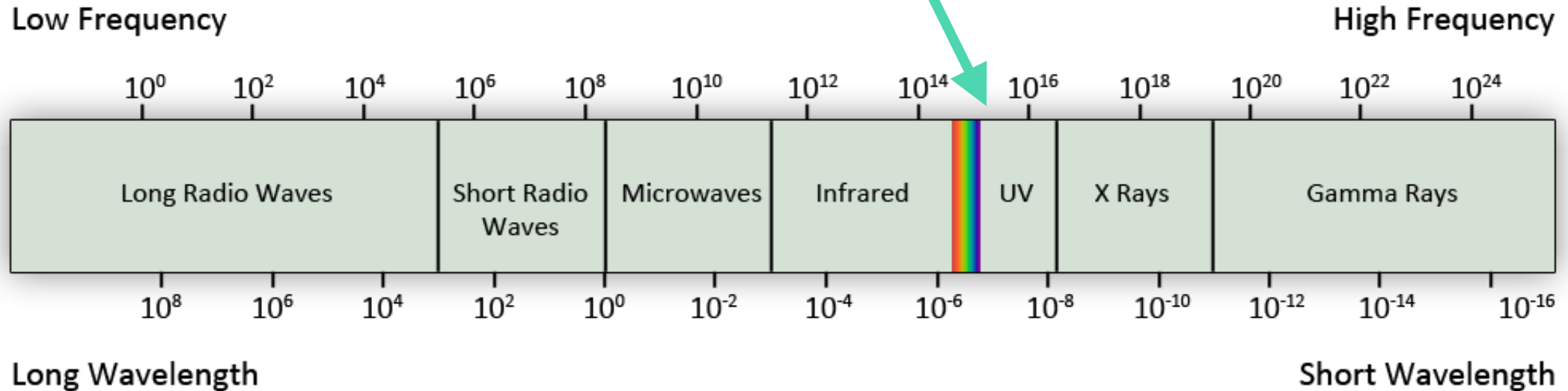
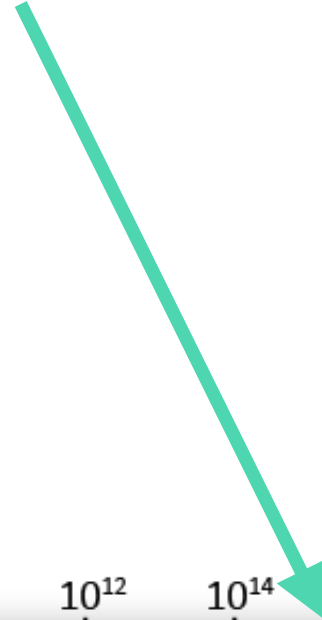
LAE galaxies

Artist representation

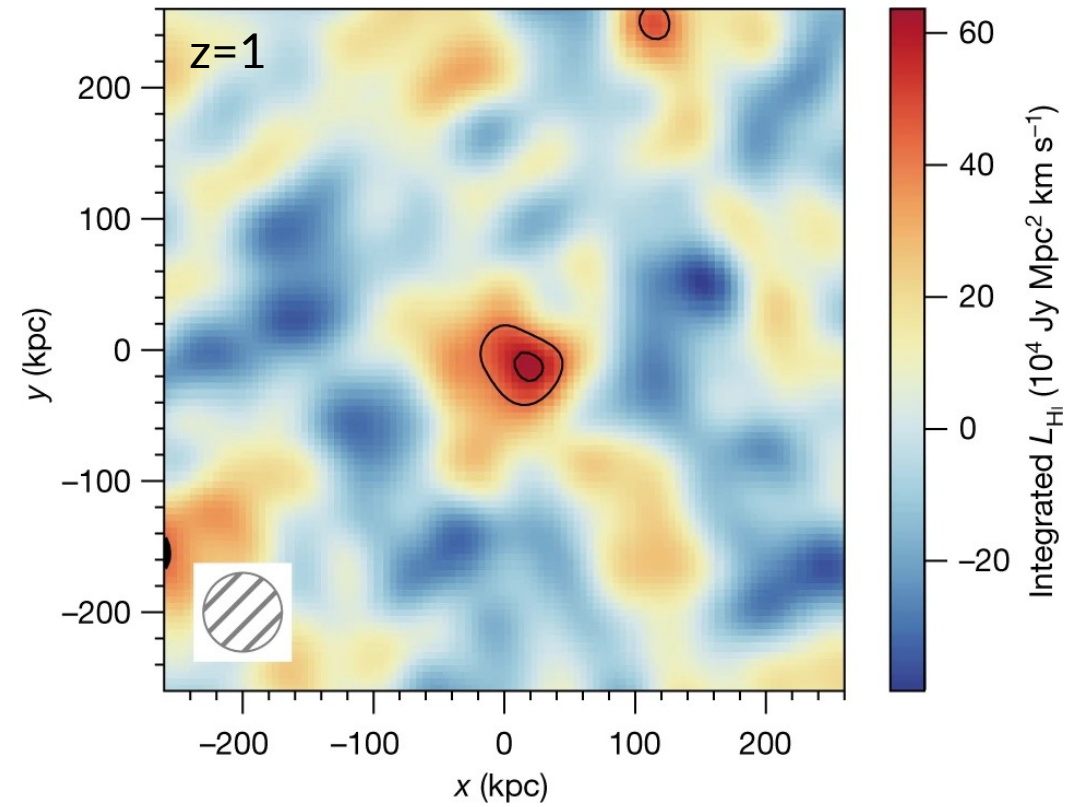
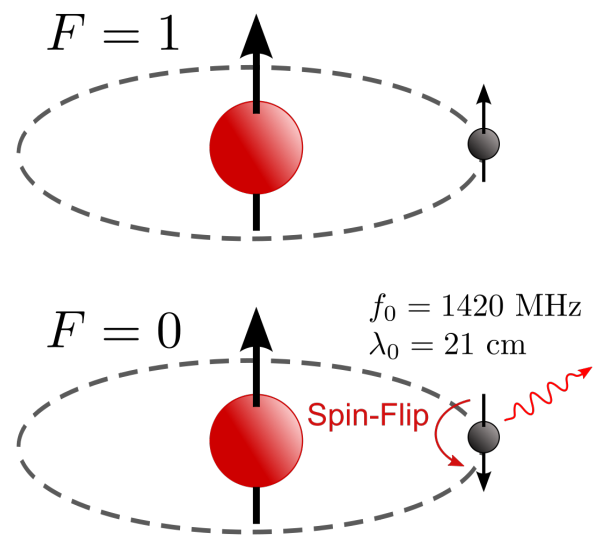
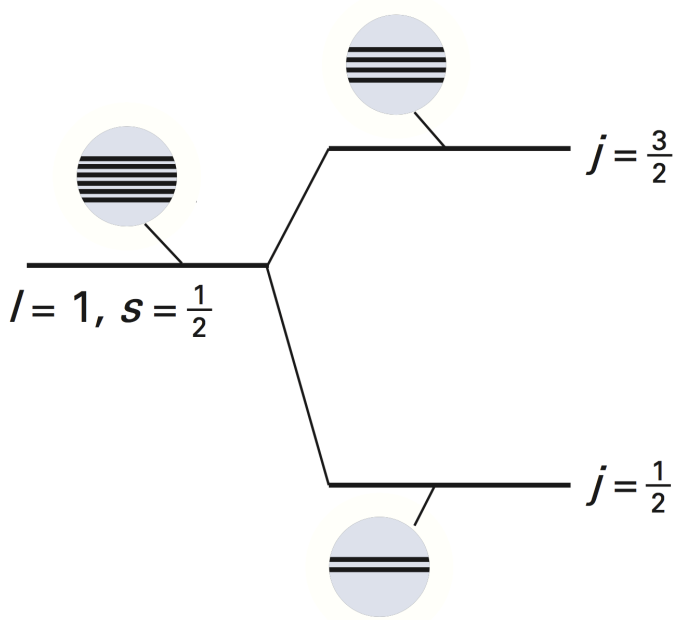
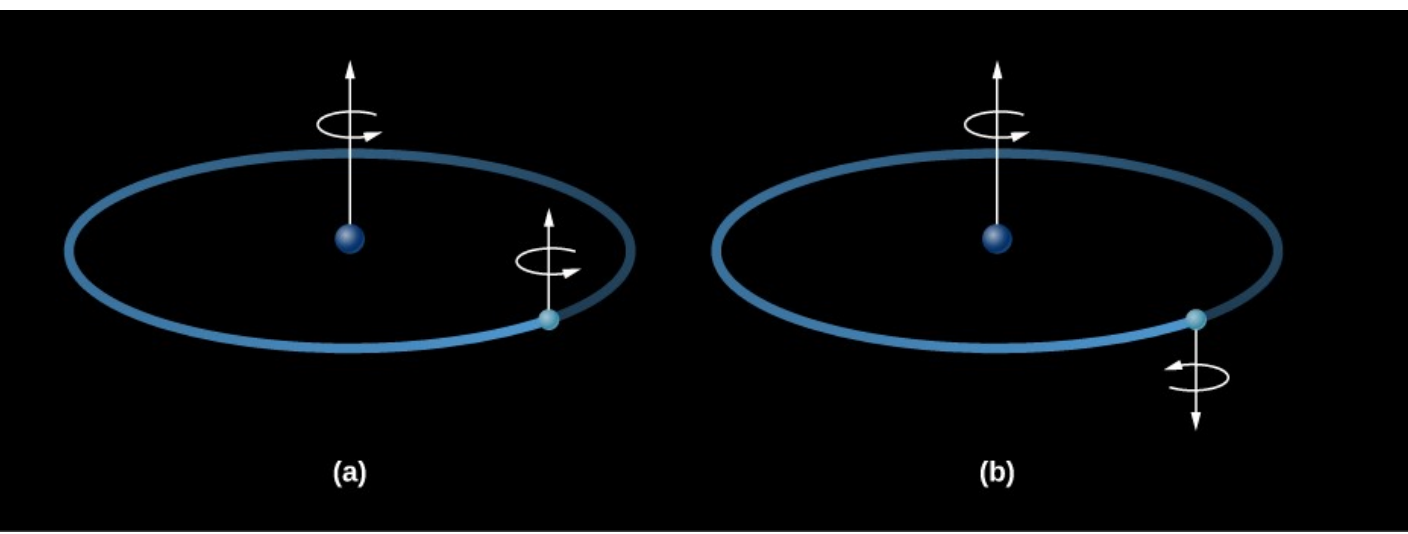




Ly-alpha observable in UV:  $2.45 \times 10^{15}$  Hz



# Spin-orbit coupling (hyperfine)



Chowdhury+, Nature, 2020

Obtained with  
Giant Metrewave Radio Telescope (PUNA)

# Multielectrons systems

- Chemists (and spectroscopists) introduced the term symbols

multiplicity  $\rightarrow 2S+1$   $\{L\}$   $\leftarrow$  orbital angular momentum  
 $J \leftarrow$  Level

## Example: Cl or NII

Six electrons

$1s^2 2s^2 2p^2$

$^3P$  ( $L=1, S=1$ ),  $^1D$  ( $L=2, S=0$ ),  $^1S$  ( $L=0, S=0$ )

Symbol	Name	Allowed Range
$L$	Total orbital angular momentum	$ l_1 + l_2 , \dots,  l_1 - l_2 $
$M_l$	Magnet Quantum number	$[m_{l1} + m_{l2}, \dots, -m_{l1} - m_{l2}]$
$M_s$	Spin Magnetic Quantum Number	$ m_{s1} + m_{s2} , \dots,  m_{s1} - m_{s2} $
$S$	Inherent Spin Number	$ s_1 + s_2 , \dots,  s_1 - s_2 $
$M$	Multiplicity	$2S+1$
$J$	Total Angular Momentum	$L + S, \dots,  L - S $

## Multielectrons:

1. Non-equivalent



Electrons with different  $n$  and/or  $l$  ( $3p^1 3d^1$ )

2. Equivalent



Electrons with same  $n$  and  $l$  ( $2p^2$ )

**Table 7.2** Terms arising from some configurations of non-equivalent and equivalent electrons

Non-equivalent electrons		Equivalent electrons	
Configuration	Terms	Configuration	Terms <sup>a</sup>
$s^1 s^1$	$1,3S$	$p^2$	$1S, 3P, 1D$
$s^1 p^1$	$1,3P$	$p^3$	$4S, 2P, 2D$
$s^1 d^1$	$1,3D$	$d^2$	$1S, 3P, 1D, 3F, 1G$
$s^1 f^1$	$1,3F$	$d^3$	$2P, 4P, 2D(2), 2F,$ $4F, 2G, 2H$
$p^1 p^1$	$1,3S, 1,3P, 1,3D$	$d^4$	$1S(2), 3P(2), 1D(2),$ $3D, 5D, 1F, 3F(2),$ $1G(2), 3G, 3H, 1I$
$p^1 d^1$	$1,3P, 1,3D, 1,3F$	$d^5$	$2S, 6S, 2P, 4P, 2D(3),$ $4D, 2F(2), 4F, 2G(2),$ $4G, 2H, 2I$
$p^1 f^1$	$1,3D, 1,3F, 1,3G$		
$d^1 d^1$	$1,3S, 1,3P, 1,3D, 1,3F, 1,3G$		
$d^1 f^1$	$1,3P, 1,3D, 1,3F, 1,3G, 1,3H$		
$f^1 f^1$	$1,3S, 1,3P, 1,3D, 1,3F, 1,3G,$ $1,3H, 1,3I$		

<sup>a</sup> The numbers in brackets indicate that a particular term occurs more than once.

**Table 7.3** Derivation of terms arising from two equivalent  $p$  electrons

Quantum number	Values															
$(m_\ell)_1$	1	1	1	1	1	1	1	1	1	0	0	0	0	0	-1	
$(m_\ell)_2$	1	0	0	0	0	-1	-1	-1	-1	0	-1	-1	-1	-1	-1	
$(m_s)_1$	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	$-\frac{1}{2}$	$-\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	$-\frac{1}{2}$	$-\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	$-\frac{1}{2}$	$-\frac{1}{2}$	$\frac{1}{2}$	
$(m_s)_2$	$-\frac{1}{2}$	$\frac{1}{2}$	$-\frac{1}{2}$	$\frac{1}{2}$	$-\frac{1}{2}$	$\frac{1}{2}$	$-\frac{1}{2}$	$\frac{1}{2}$	$-\frac{1}{2}$	$-\frac{1}{2}$	$\frac{1}{2}$	$-\frac{1}{2}$	$\frac{1}{2}$	$-\frac{1}{2}$	$-\frac{1}{2}$	
$M_L = \sum_i (m_\ell)_i$	2	1	1	1	1	0	0	0	0	0	-1	-1	-1	-1	-2	
$M_S = \sum_i (m_s)_i$	0	1	0	0	-1	1	0	0	-1	0	1	0	0	-1	0	
Pairs of values of $M_L$ and $M_S$ can be rearranged as follows:																
$M_L$	2	1	0	-1	-2	1	0	-1	1	0	-1	1	0	-1	0	
$M_S$	<u>0 0 0 0 0</u>					<u>1 1 1 0 0</u>					<u>0 0 -1 -1 -1</u>					0
	${}^1D$					${}^3P$					${}^1S$					

# Hund's rules

---

- The state with the largest value of  $S$  is the most stable
- For states with the same  $S$ , the state with largest  $L$  is the most stable
- If the states have same  $S$  and  $L$ , the state with the smallest  $J$  is the most stable (if half filled)
- If more than half filled the state with largest  $J$  is the most stable

# Term symbols: more

- Degeneracy factor  $g = (2S + 1) \times (2L + 1)$
- Example:  $^3P$  term ( $S = 1, L = 1$ ) has a  $g = 3 \times 3 = 9$
- If we take into consideration spin-orbit coupling these states are split into different fine structure levels
- We have then degeneracy  $g = 2J + 1$

$\text{multiplicity} \rightarrow 2S + 1$   $\{L\}$   $\leftarrow$  orbital angular momentum  
 $J \leftarrow \text{Level}$



# Selection rules

Are restrictions that govern the possible (or allowed) transitions from a state to another

$$\Delta L = \pm 1$$

**Follow conservation of angular momentum**

$$\Delta S = 0$$

**Spin can only change by 0**

$$\Delta J = 0, \pm 1$$

**Follow conservation of angular momentum**

# Selection rules: Spin

**Singlet-Singlet, Doublet-Doublet etc. allowed**

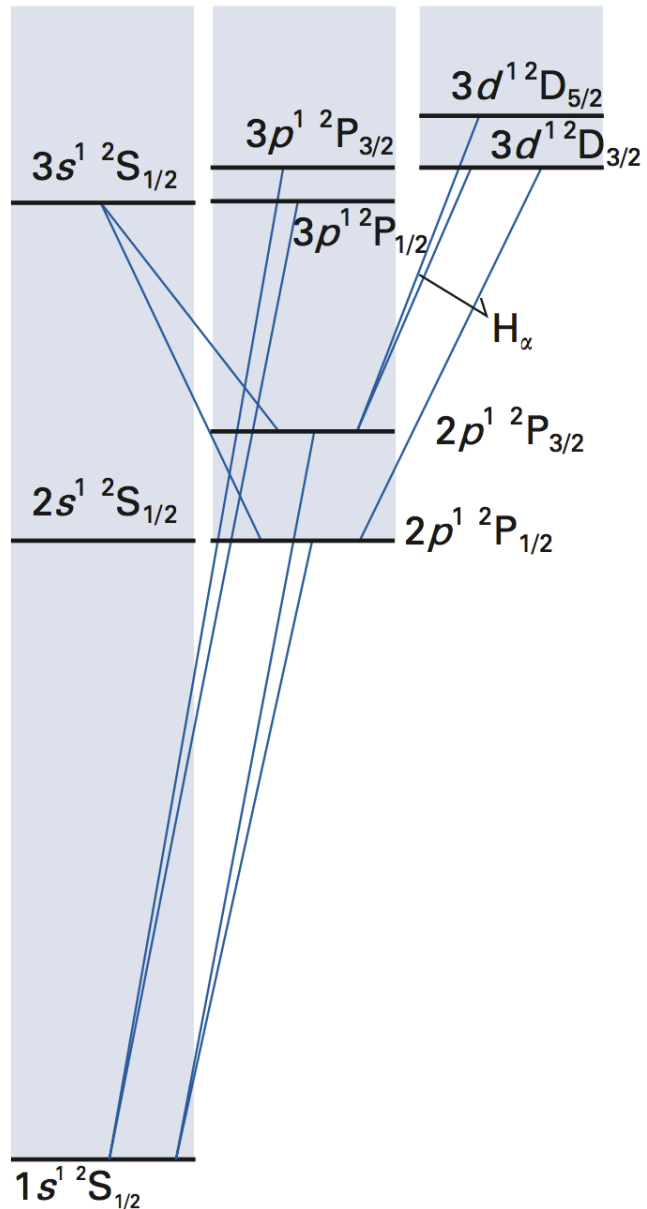
**Singlet-Triplet etc. forbidden**

**It is also a consequence of the Pauli principle**

**allowed:  $s \leftrightarrow p, p \leftrightarrow d$**

**forbidden:  $d \leftrightarrow d, p \leftrightarrow p$**

# Selection rules Lyman series of H



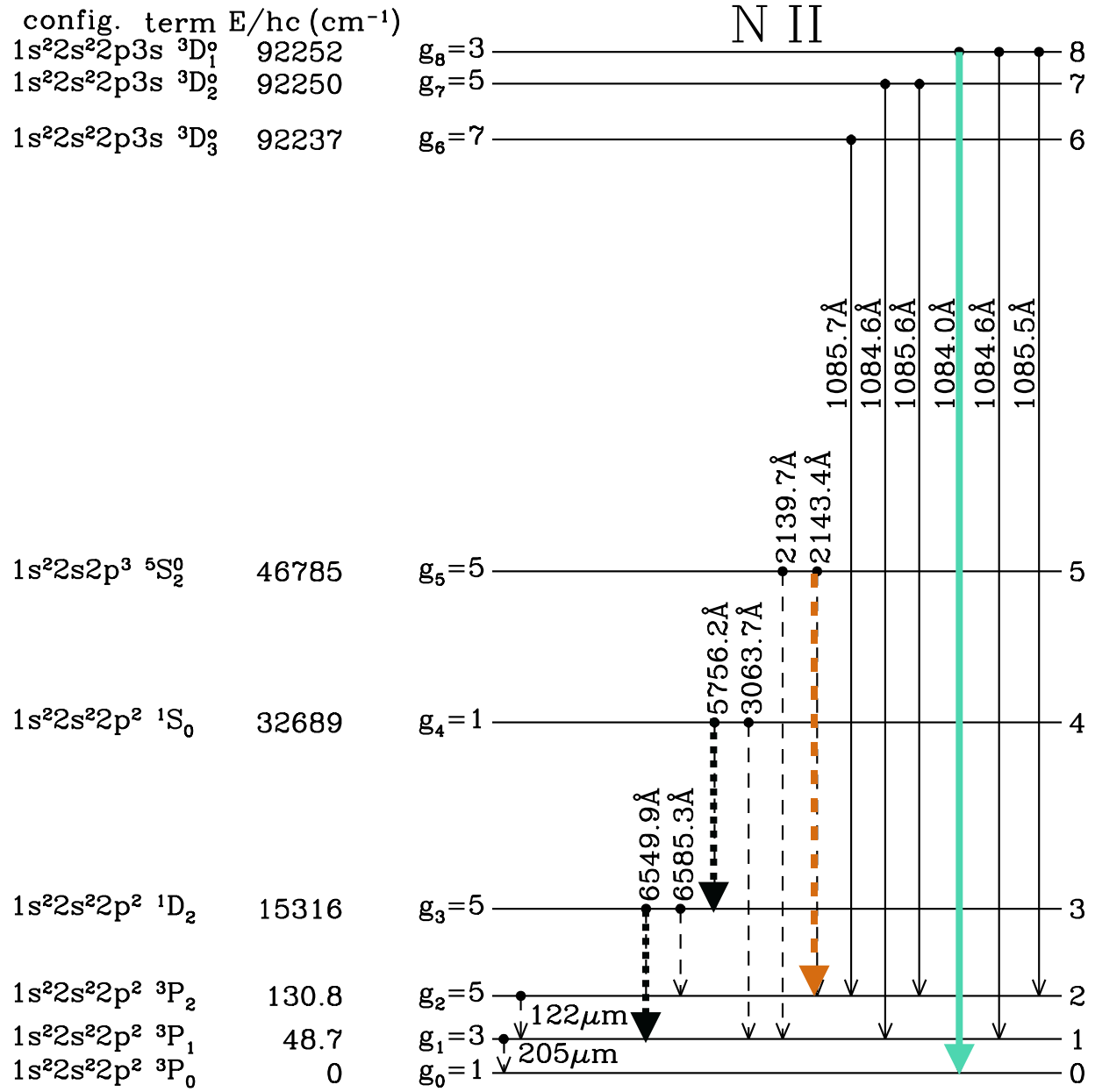
$$np \ ^2P_{1/2} \rightarrow 1s \ ^2S_{1/2} \quad \begin{pmatrix} \Delta L = 1 \\ \Delta S = 0 \\ \Delta J = 0 \end{pmatrix}$$

$$np \ ^2P_{3/2} \rightarrow 1s \ ^2S_{1/2} \quad \begin{pmatrix} \Delta L = 1 \\ \Delta S = 0 \\ \Delta J = -1 \end{pmatrix}$$

$$\Delta L = \pm 1$$

$$\Delta S = 0$$

$$\Delta J = 0, \pm 1$$



# Allowed transition

<sup>3</sup>D<sub>1</sub> - <sup>3</sup>P<sub>0</sub>

$A_{ul} = 2.18 \times 10^8 \text{ s}^{-1}$

# Spin forbidden transition

<sup>5</sup>S<sub>2</sub> - <sup>3</sup>P<sub>2</sub>

$A_{ul} = 1.27 \times 10^2 \text{ s}^{-1}$

# Forbidden transition

<sup>1</sup>D<sub>2</sub> - <sup>3</sup>P<sub>1</sub>

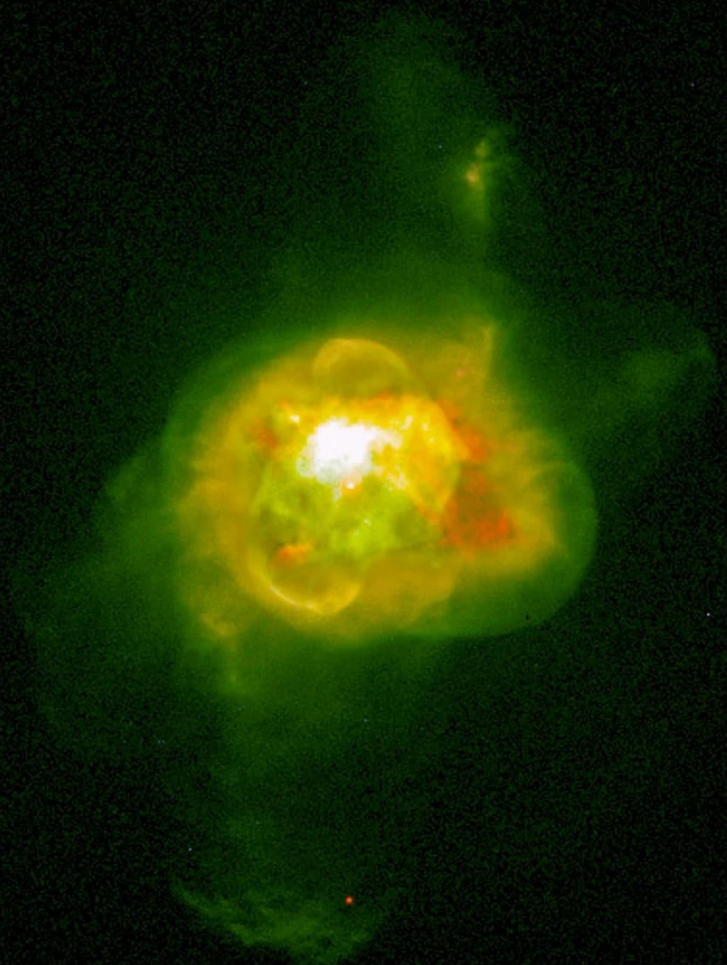
$A_{ul} = 1.17 \times 10^0 \text{ s}^{-1}$

# General rules

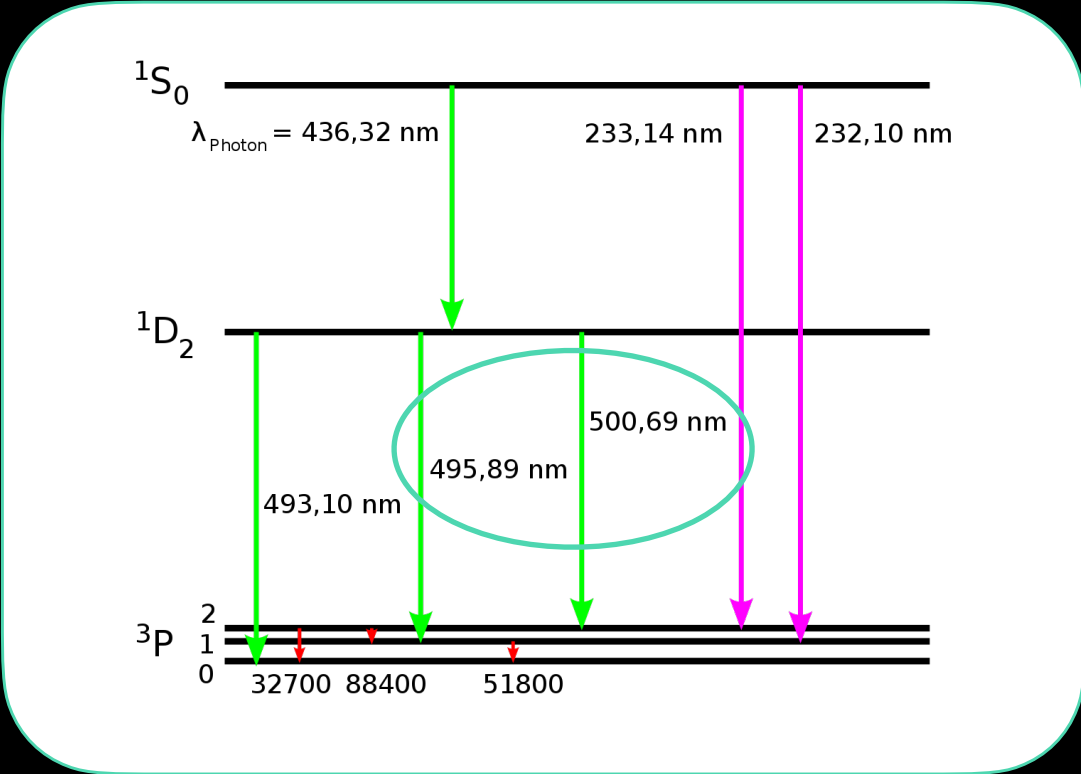
---

- Intersystem lines are  $10^6$  times weaker than permitted transitions
- Forbidden transitions are  $10^2$ - $10^6$  times weaker than intersystem ones
- **Forbidden transitions** are important in astrophysics in particular in the low-dense environment where timescales are much longer and slow decay could occur (collisions very rare)

Strongest forbidden lines observed in the ISM are from doubly ionized oxygen [OIII]

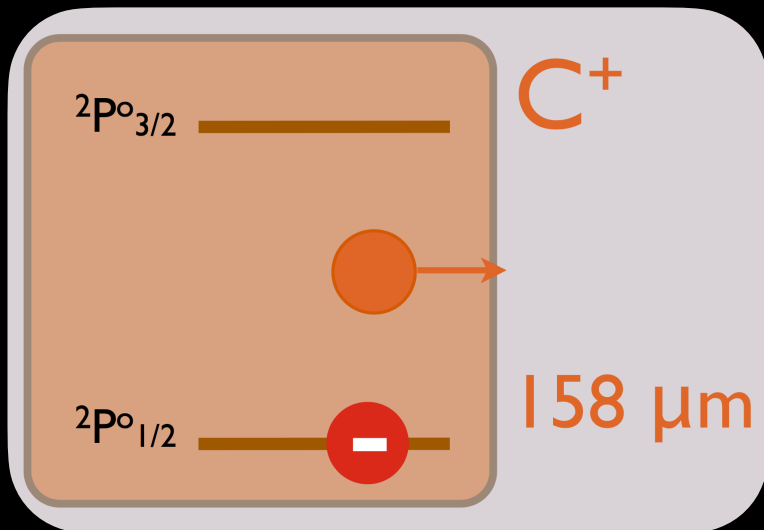


NGC6210

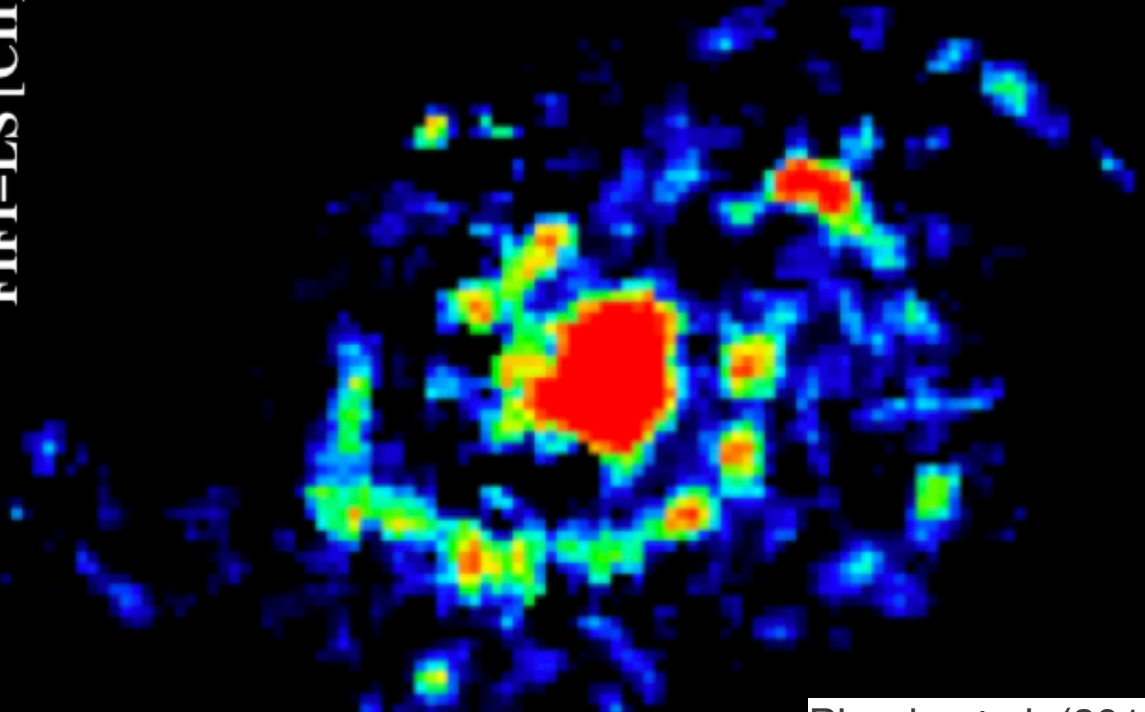




# 158 micron [CII] transition

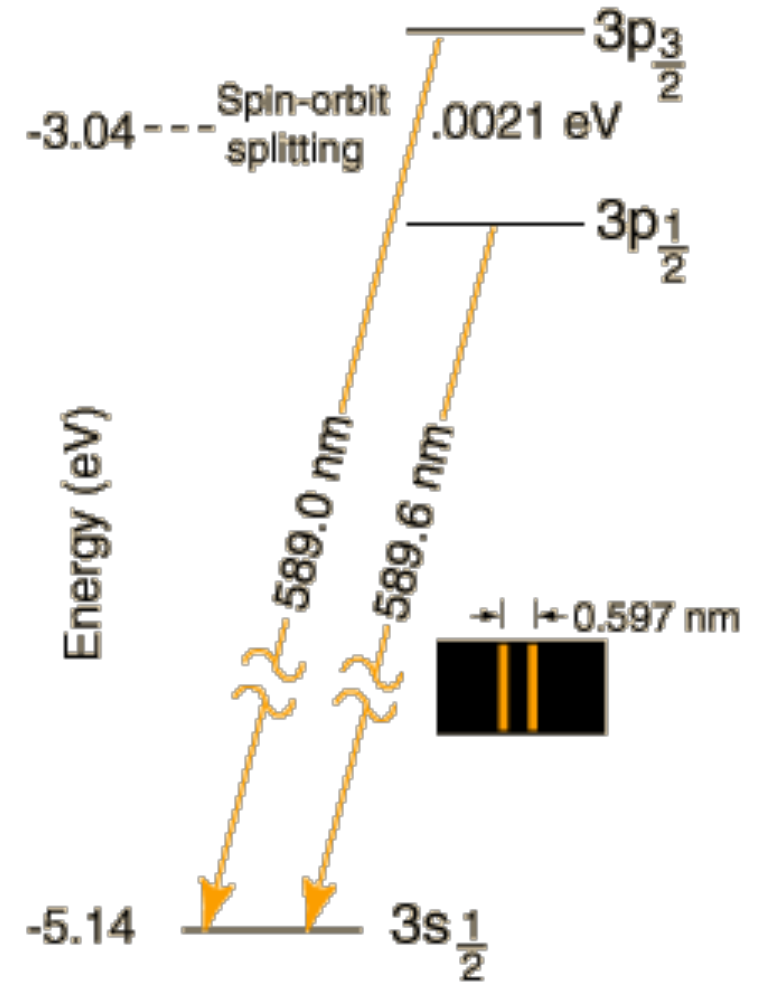
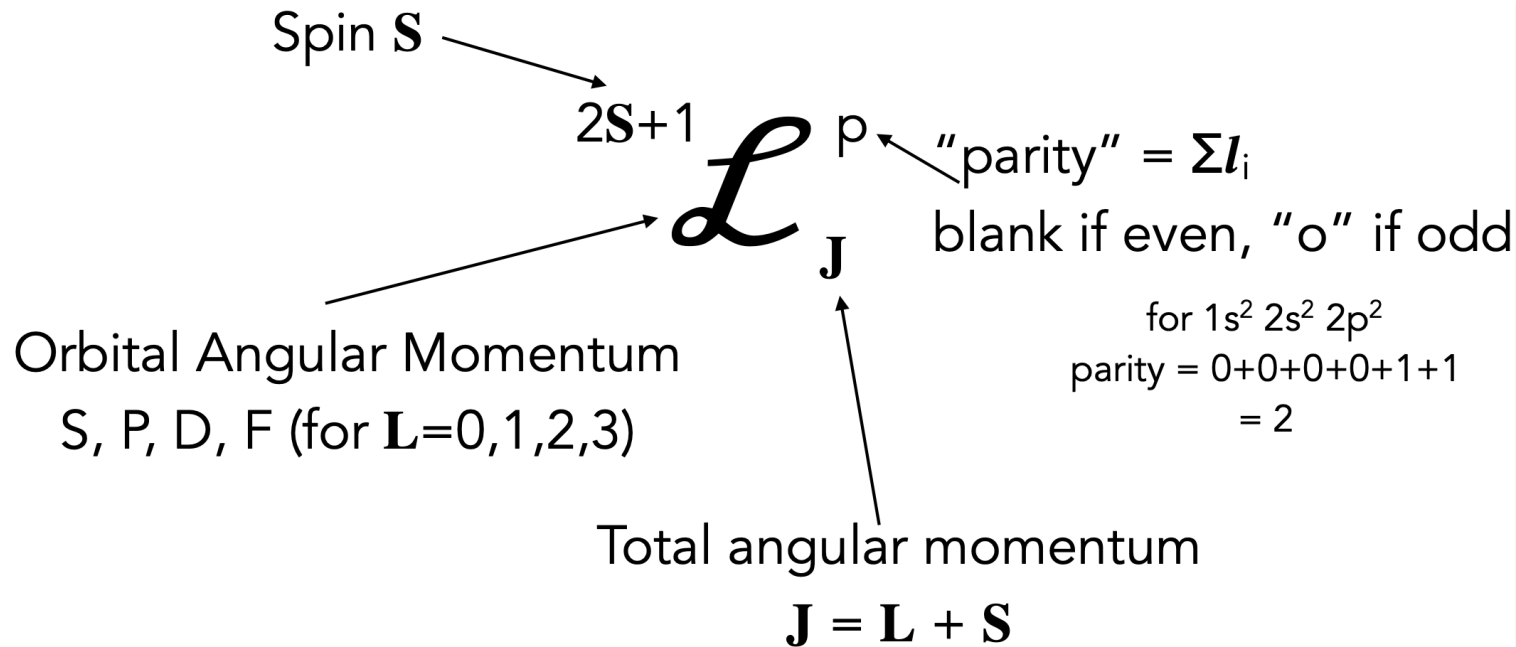


FIFI-LS [CII]



# Term symbols: Na doublet (exercise)

Sodium, Na, Z=11





# Most of this information

**Molecular Quantum Mechanics (Atkins & Friedman)**

**Modern Spectroscopy (Hollas)**